

Optics

Planck's radiation

Consider a cavity of length L with metallic walls. Since metals are good conductors of electricity, the electric field at each wall will have to be zero. Then the only stationary modes allowed are:

$$\lambda = \frac{2L}{n}, \text{ where } n \text{ is a positive integer.}$$

That mode will have an energy equal to $E_n = Nhf = N \frac{hc}{\lambda} = N \frac{hcn}{2L}$, where N is any nonnegative integer.

According to statistical physics the average energy stored in this mode will be:

$$\langle E_n \rangle = \frac{\sum_{N=0}^{\infty} N \frac{hcn}{2L} \exp\left(-N \frac{hcn}{2Lk_B T}\right)}{\sum_{N=0}^{\infty} \exp\left(-N \frac{hcn}{2Lk_B T}\right)} = \frac{hcn}{2L \left[\exp\left(\frac{hcn}{2Lk_B T}\right) - 1 \right]}$$

If we wanted to calculate the density of energy we would need to add the contribution from each mode and divide by the volume of the cavity. Let's assume the cavity is a cube, so there will be a contribution from each possible direction and each polarization (a factor of 2).

The modes will be combinations of three different orientations: (n_x, n_y, n_z) with an energy:

$$E_{n_x, n_y, n_z} = \frac{2hcn}{2L \left[\exp\left(\frac{hcn}{2Lk_B T}\right) - 1 \right]}, \text{ where } n = \sqrt{n_x^2 + n_y^2 + n_z^2}$$

Now to find out the number of modes in a wavelength interval notice that the number of modes grows like 1/8 the volume of a sphere:

$$\# \text{ modes } (n < n_o) = \frac{1}{8} \left(\frac{4}{3} \pi n_o^3 \right) \rightarrow \# \text{ modes } (\lambda > \lambda_o) = \frac{1}{8} \left(\frac{4}{3} \pi \left[\frac{2L}{\lambda_o} \right]^3 \right) = \frac{4}{3} \pi \frac{L^3}{\lambda_o^3}$$

Taking a derivative to find the density of modes:

$$\frac{d(\# \text{ modes})}{d\lambda} = -4\pi \frac{L^3}{\lambda_o^4}, \text{ where the sign indicates that the larger the wavelength the less the number of modes.}$$

Going back to the energy, we can now calculate the density of energy per wavelength and per volume of the cavity:

$$U(\lambda) = \frac{4\pi L^3}{\lambda^4} \frac{2hc\nu}{2L \left[\exp\left(\frac{hcn}{2Lk_B T}\right) - 1 \right] L^3}$$

$$U(\lambda) = \frac{8\pi hc}{\lambda^5 \left[\exp\left(\frac{hc}{k_B T \lambda}\right) - 1 \right]}$$

To get the radiation we multiply by $c/4$, the speed of light and a factor of four due to the geometry of the vector distribution.

$$R = \frac{2\pi hc^2}{\lambda^5} \frac{1}{e^{\frac{hc}{\lambda k_B T}} - 1}$$